



# Development of a real-time tool for deriving the optical properties of aerosol particles in PANGEA observatory, using ground-based passive remote sensing measurements

Thesis Supervisor

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**Master thesis**

18 September 2020, Athens



- ▣ Thesis Scope
- ▣ Aerosol research
- ▣ PANhellenic GEophysical observatory of Antikythera (PANGEA)
- ▣ AErosol RObotic NETwork (AERONET)
- ▣ Developed tool and implementation
- ▣ Results
- ▣ Future work
- ▣ Summary

# Thesis Scope



**Aerosols** can affect Earth's energy budget, by directly scattering, absorbing and emitting the incoming solar radiation, and by indirectly influencing the cloud properties.

The atmospheric aerosols play an important role to:

- ▣ Weather
- ▣ Air quality
- ▣ Climate
- ▣ Human health
- ▣ Ecosystems

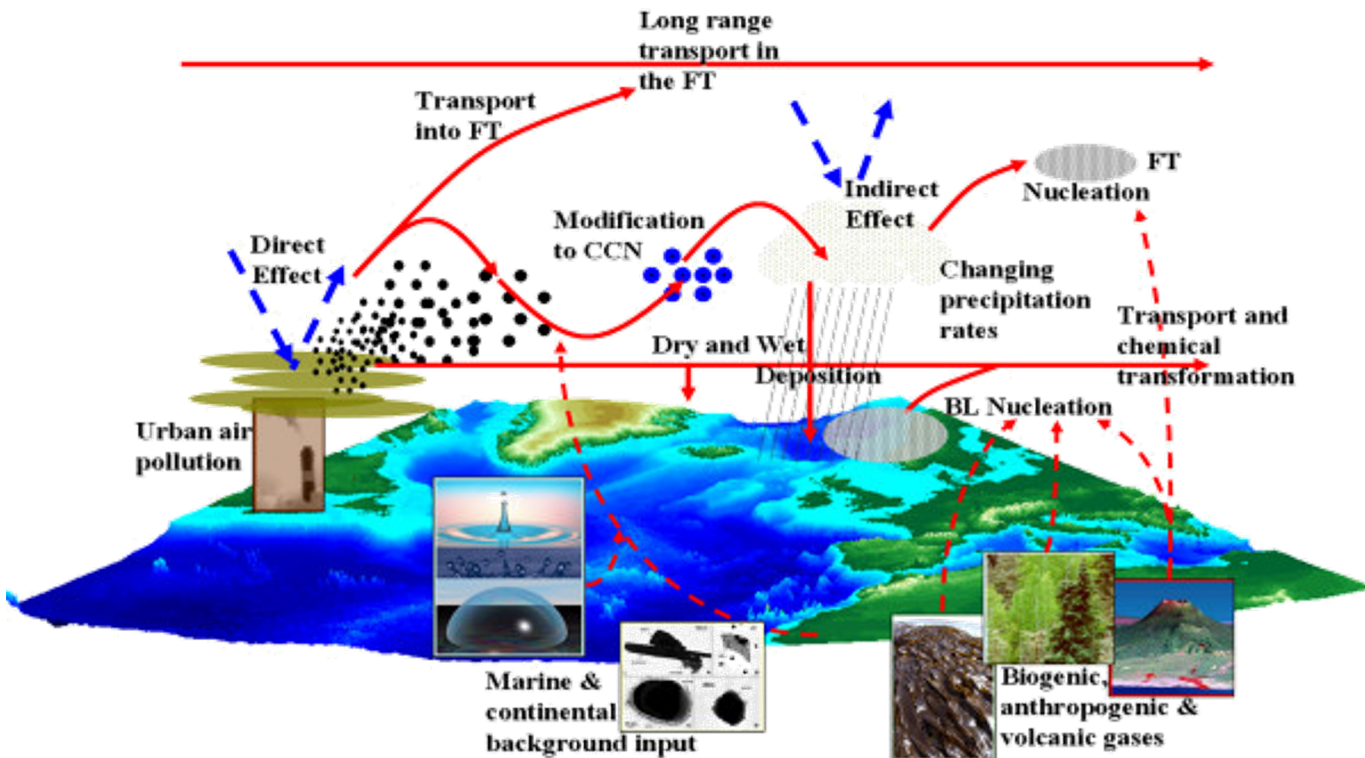


Diagram of the life cycle of aerosol in the atmosphere, showing emission, deposition and transport processes and the action of aerosol while in the atmosphere.

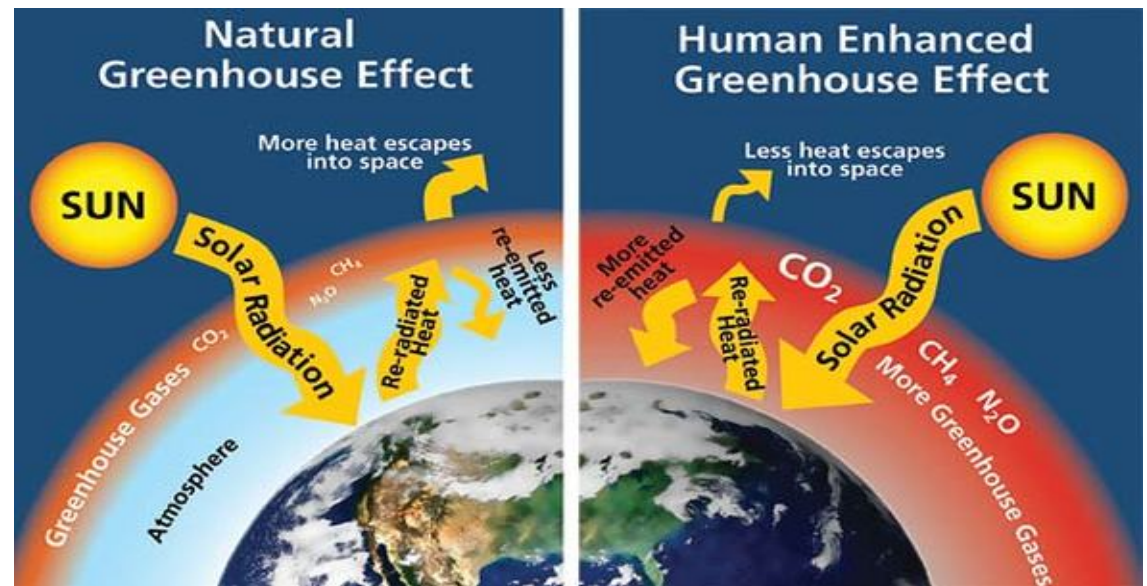
<http://www.cas.manchester.ac.uk/resactivities/aerosol/>



# Importance of this work



- ▣ This work calculates the optical properties of aerosol particles.
- ▣ The **importance** of aerosol optical properties for climate studies.
  - a) Because aerosol impact the Earth-atmosphere budget
  - b) Their chemical reaction affect the weather and can harm human health





**Tool description:** Development of a tool for deriving the optical properties of aerosol particles in PANGEA observatory at near-real-time, using ground-based passive remote sensing measurements.

Development steps:

- **Create** a tool that automatically downloads the sun-photometer measurements from the Antikythera AERONET station (at PANGEA), in near real-time.
- **Display** them in graphical plots.
- Use the **Mie scattering code** to calculate the particle optical properties: the extinction, absorption, scattering and backscatter coefficients.

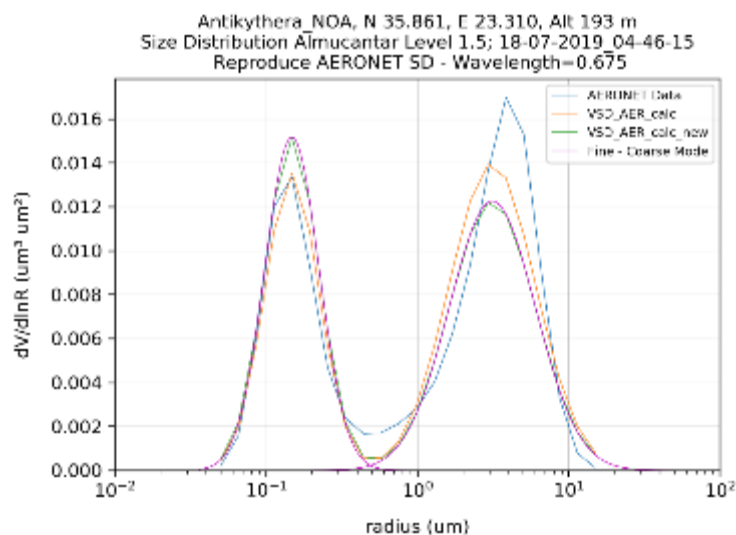


## ■ Create a new tool that:

■ **Automatically downloads** the AERONET sun-photometer products (size and refractive index of aerosol particles) derived from measurements at PANGEA.

■ **Display** them in graphical plots in near real-time.

■ **Calculate** the optical properties of the particles, when they have spherical shapes.



```
ext_all=8.326684453680716e-05 sca_all=7.411795168585089e-05  
abs_all=9.14889285095627e-06 bac_all=1.7400255234561573e-06  
ext_f=7.183912358268646e-05 sca_f=6.622154404926358e-05  
abs_f=5.617579533422876e-06 bac_f=1.4251190852857044e-06  
ext_c=1.1427720954120696e-05 sca_c=7.896407636587302e-06  
abs_c=3.531313317533394e-06 bac_c=3.1490643817045284e-07
```



- It is a **near real-time tool**

The tool can calculate in few seconds the optical properties of aerosols after the data are published in the AERONET





**PAN**hellenic **GE**ophysical  
observatory of **Antikythera**



PANGEA observatory in Antikythera island has been installed by the **National Observatory of Athens** (NOA) in order to respond to the scientific need for accurate climatic observations in the Mediterranean.

## Why **Antikythera** ?

- ▣ crossroad of Europe, Africa and Asia
- ▣ anthropogenic activities are very few, due to its population
- ▣ ideal place for monitoring natural aerosols
- ▣ transported anthropogenic aerosols from big cities in Greece and the surrounding Mediterranean countries



*Antikythera NOA equipment  
(from National Documentation Centre, [www.ekt.gr](http://www.ekt.gr))*

## Available instruments in PANGEA:

- ▣ **PolyXT lidar** (EARLINET)
- ▣ **CIMEL sun-photometer** (AERONET)



**What** does CIMEL Sun-Photometer do?

▣ **Provides:**

- ▣ AOD from **direct-sun measurements**
- ▣ Microphysical properties of the aerosol particles from **diffused light measurements**

**How** does CIMEL Sun-Photometer **acquire** measurements?

- ▣ Every day between sunlight and dawn
- ▣ In **nine spectral bands:**
  - ▣ 340, 380, 440, 500, 670, 870, 940, 1020 and 1640 nm.

**The CIMEL sun-photometer products provide us with the microphysical properties the developed tool uses for the calculation of the aerosol optical properties.**

It's a **passive remote sensing** instrument



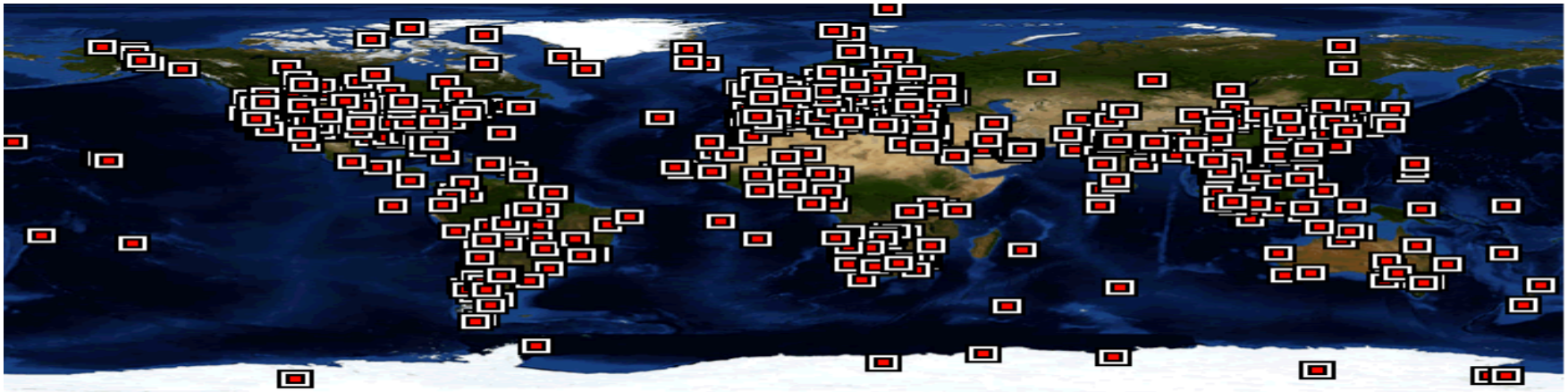
*CIMEL sun-photometer  
(from <https://aeronet.gsfc.nasa.gov>)*

# **A**Erosol **R**Obotic **N**ETwork



## What is the AERONET?

- It is a network of **ground-based** sun photometers



*The worldwide spatial distribution of the AERONET stations*

- It provides observations of the **optical** and **microphysical properties** of the atmospheric aerosols

## How the AERONET data products can be accessed ?

- Through the **AERONET web download tool** in AERONET site <https://aeronet.gsfc.nasa.gov/>
- Through the **AERONET web services**

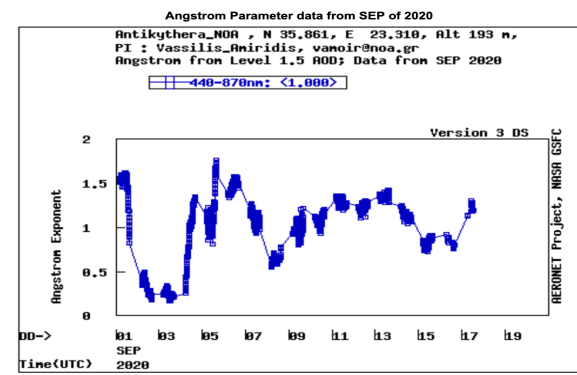
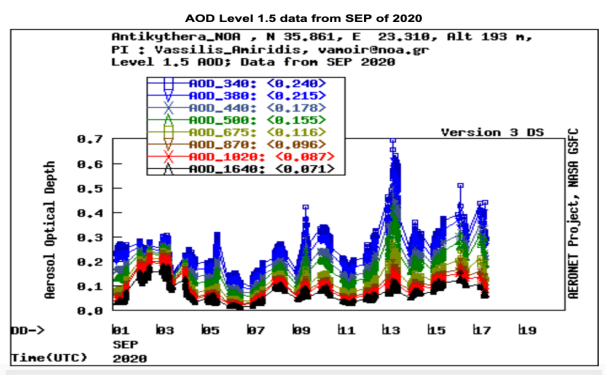


# AERONET...Aerosol products

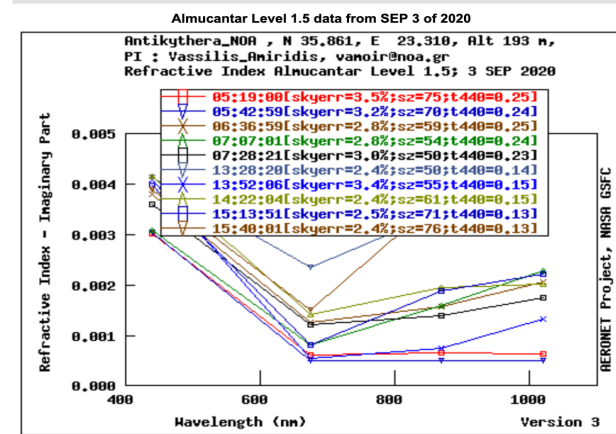
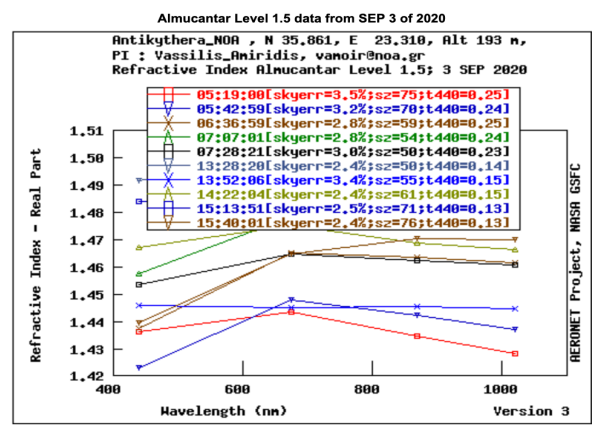
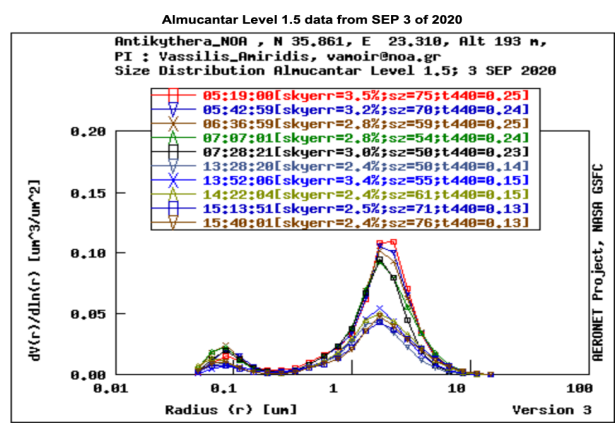


AERONET data display uses a **user-friendly interface** that includes **graphical representations** to display the optical and microphysical properties of the aerosol particles.

## AERONET DATA PRODUCTS



## AERONET INVERSION DATA PRODUCTS



Developed tool and  
implementation

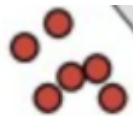


## What is our aim?

- ▣ To calculate the **optical properties of spherical aerosol particles** in PANGEA observatory based on the AERONET measurements
  - The **scattering** coefficient
  - The **backscatter** coefficient
  - The **absorption** coefficient
  - The **extinction** coefficient



for **fine** (diameter  $<2.5 \mu\text{m}$  ) and **coarse** (diameter  $>2.5 \mu\text{m}$  ) mode.



Fine



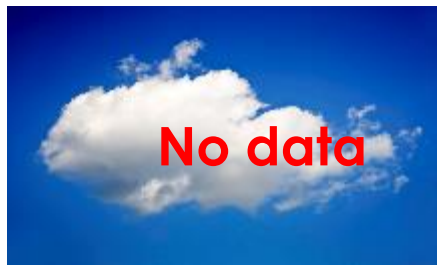
Coarse





- ▣ **What** data are used from Antikythera\_NOA AERONET station?
  - ▣ The **microphysical properties** of the aerosols
    - The **volume size distribution** ( $dv/d\ln r$ )
    - The **complex refractive index**
    - The **sphericity percentage**
  
- ▣ **How often the station gives data?**

if there are clouds





## What data are used from Antikythera\_NOA AERONET station?

- ▣ The **microphysical properties** of the aerosols
  - The **volume size distribution** ( $dv/d\ln r$ )
  - The **refractive index**
  - The **sphericity percentage**

## How often the station gives data?

Otherwise



UTC /Local time	UTC /Local time
~ 5:00/8:00	~ 13:34/16:34
~ 5:25/8:25	~ 14:28/17:28
~ 6:13/9:13	~ 14:58/17:58
~ 6:43/9:43	~ 15:49/18:49
~ 7:34/10:34	~ 16:13/19:13
~ 8:25/11:25	No data
~ 12:25/15:25	Until next day

After how many minutes we can **access** the **data**?

**~25'**



Loading...



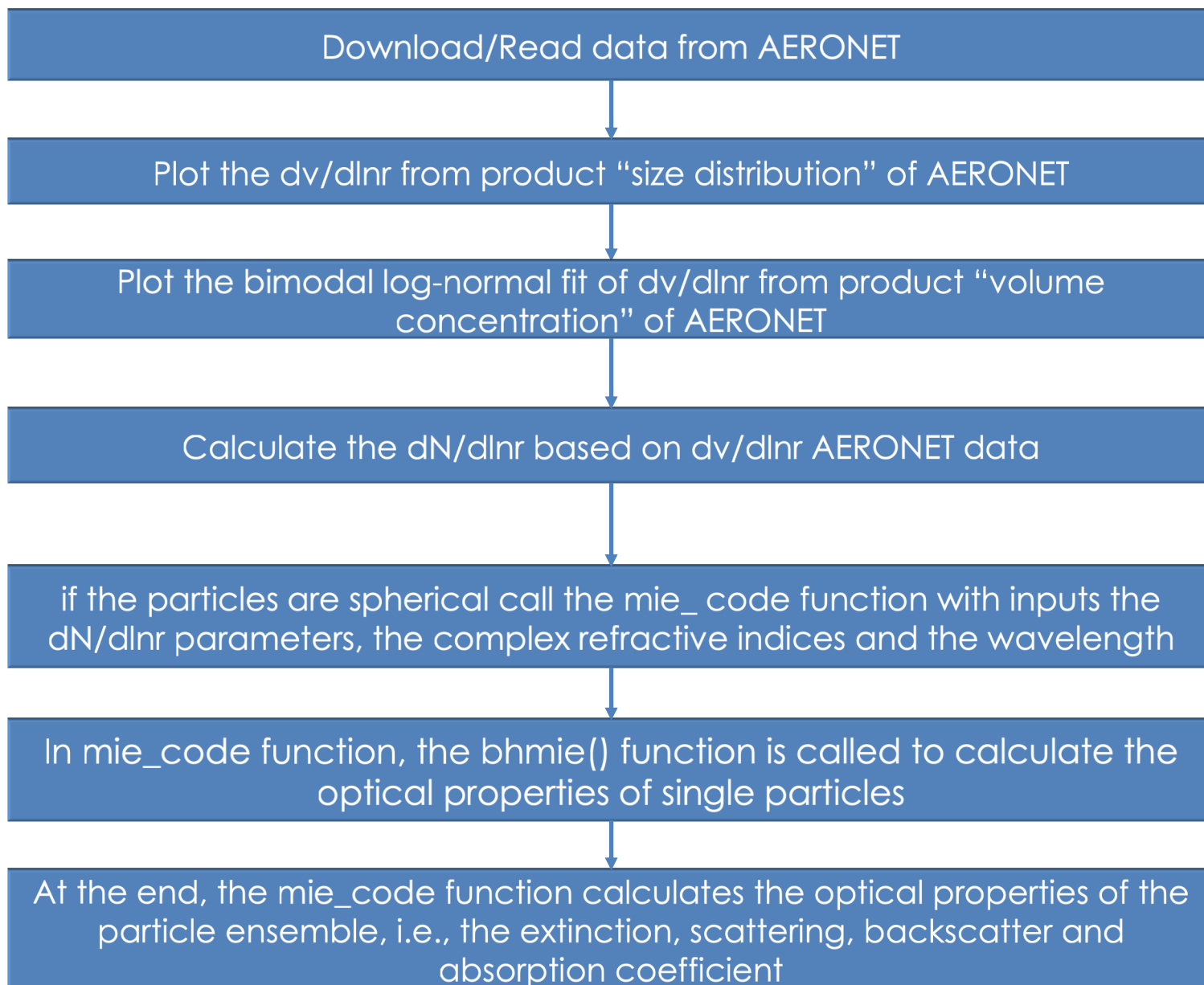
- ▣ Programming language: **Python version 3.7**



- Open Source
- Object Oriented
- Big Supported Community
- Easy to learn

- ▣ Scientific Python development environment: **Spyder with Anaconda**







# Code analytic detailed description



Check arguments  
(start/end date, wavelength)

The **user** specifies the **wavelength** and the date or the **period** that wants to study.

By **default**, the **wavelength** is equal to **0.44  $\mu\text{m}$**  and the period is the current day.

Web request to Antikythera  
AERONET server  
for AERONET data/web service

Check if  
there are  
data for  
these  
dates

**NO**

The system **exits** by informing the user that for these dates there are no data.

**YES**

The program **downloads** automatically the relevant **data**.



- In the **case that there are data**, there is an **iteration**, because **there can be more than one measurements (at different times) in the day**.

## AERONET product

## Parameters

Download data



Read data

**Product.asy** → **Sphericity\_Factor (%)**

**Product.rin** → **Refractive\_Index  
Real\_Part & Imaginary\_Part**  
(440 nm, 678 nm, 870 nm, 1020 nm)

**Product.siz** → **Size distribution at 22 size bins**

**Product.vol** → **Parameters of the bimodal log-normal fit of the 22-size-bin volume size distribution:** the geometric standard deviation, total volume concentration, geometric mean volume radius, for both fine and coarse modes



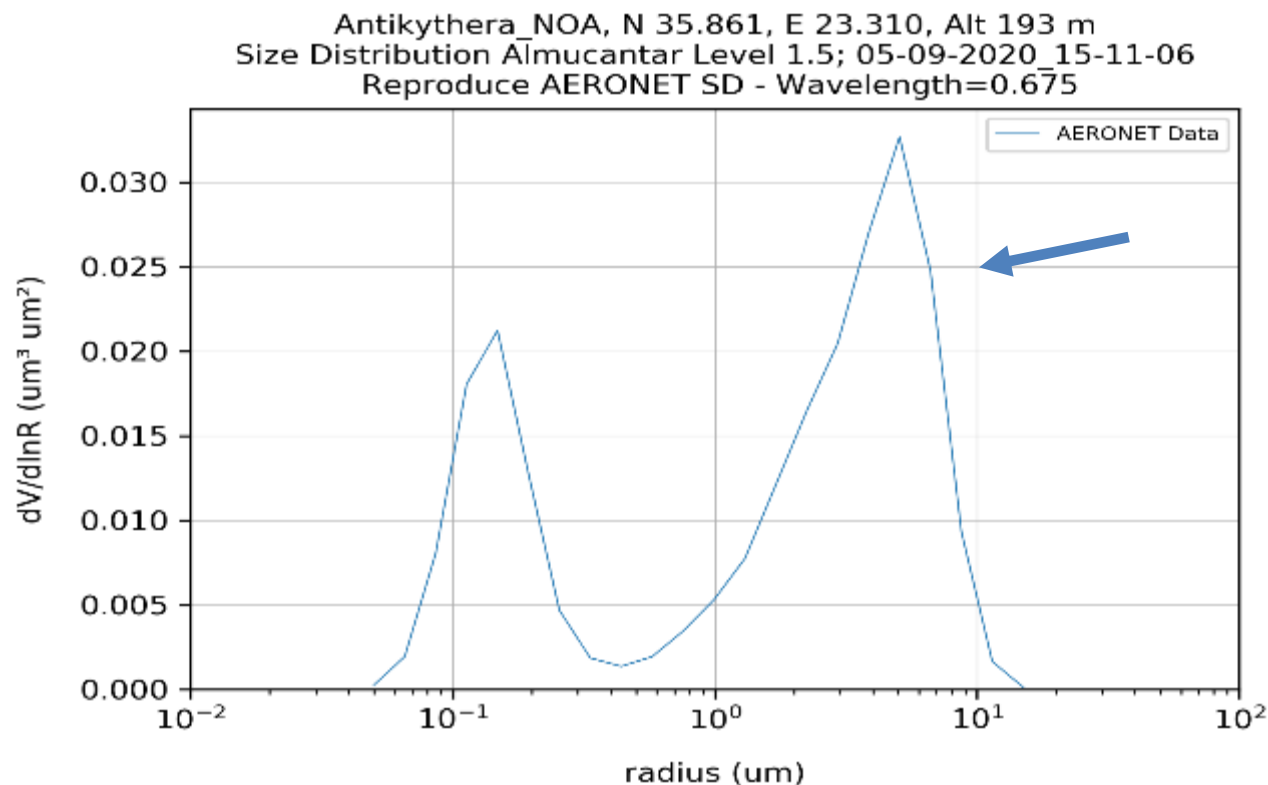
Download data

Read data

Do the  
Calculations &  
Create the plots

## 1<sup>st</sup> step

The **blue line** represents the **22-size-bin Volume Size Distribution** ( $dv/d\ln R$ ) acquired from the AERONET product (product.siz).





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Do the  
Calculations &  
Create the plots

## 2<sup>nd</sup> step

Based on **the (six) parameters** of the bimodal log-normal fit of the 22-size-bin volume size distribution, we calculate the  $dv/d\ln r$  fit:

$$\frac{dV}{d\ln r} = \frac{volc_f}{\sqrt{2\pi} \ln(std_f)} \exp\left(-\frac{(\ln(r) - \ln(vmr_f))^2}{2\ln(std_f)^2}\right) + \frac{volc_c}{\sqrt{2\pi} \ln(std_c)} \exp\left(-\frac{(\ln(r) - \ln(vmr_c))^2}{2\ln(std_c)^2}\right)$$

where  $\frac{dV}{d\ln r}$ : the particle volume size distribution

$volc_f$ : the total volume concentration of fine mode

$volc_c$ : the total volume concentration of coarse mode

$vmr_f$ : volume geometric mean radius of fine mode

$vmr_c$ : volume geometric mean radius of coarse mode

$std_f$ : geometric standard deviation of fine mode

$std_c$ : geometric standard deviation of coarse mode





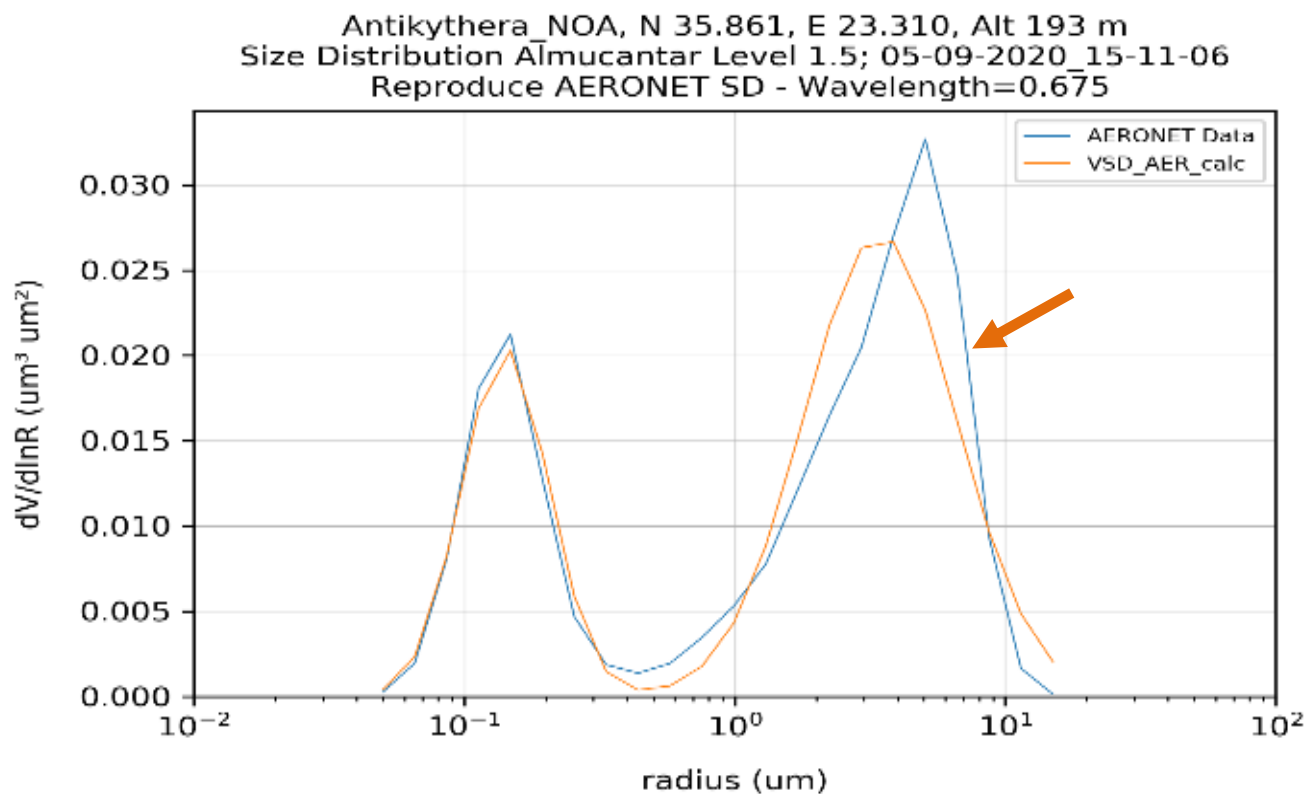
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Do the  
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Create the plots

## 2<sup>nd</sup> step

The **orange line** is the **bimodal log-normal fit of the 22-size-bin Volume Size Distribution (dv/dlnr\_fit)**. It fits the blue line using only **six parameters** (volc\_f, volc\_c, vmr\_f, vmr\_c, std\_f, std\_c).





Download data



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Do the  
Calculations &  
Create the plots

## 2<sup>nd</sup> step

### Why we need this step?

Because the **Mie** scattering **code** has as **input** a **bimodal lognormal size distribution**.

This size distribution is described based on these six parameters.

### Parameters

- **Total volume concentration** (volc\_f, volc\_c)
- **Volume geometric mean radius** (vmr\_f, vmr\_c)
- **Geometric standard deviation** (std\_f, std\_c)



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Do the  
Calculations &  
Create the plots

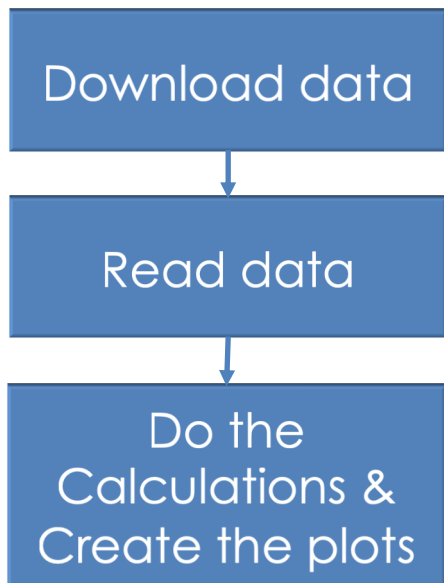
2<sup>nd</sup> step

More specifically

The **input** of the **Mie** scattering **code** is a **bimodal lognormal number size distribution**.

And how can I calculate it?

$$\frac{dN}{d\ln r} = \frac{Nc_f}{\sqrt{2\pi} \ln(std_f)} \exp\left(-\frac{(\ln(r) - \ln(nmr_f))^2}{2\ln(std_f)^2}\right) + \frac{Nc_c}{\sqrt{2\pi} \ln(std_c)} \exp\left(-\frac{(\ln(r) - \ln(nmr_c))^2}{2\ln(std_c)^2}\right)$$



## 2<sup>nd</sup> step

In order to **calculate the six parameters of the number size distribution** ( $N_{c\_f}$ ,  $N_{c\_c}$ ,  $n_{mr\_f}$ ,  $n_{mr\_c}$ ,  $std\_f$ ,  $std\_c$ ) **based on** the six parameters of the **volume size distribution** ( $volc\_f$ ,  $volc\_c$ ,  $vmr\_f$ ,  $vmr\_c$ ,  $std\_f$ ,  $std\_c$ ) we use these formulas:

$$\frac{dN}{dnr} = \frac{N_{c\_f}}{\sqrt{2\pi} \ln(std\_f)} \exp\left(-\frac{(\ln(r) - \ln(nmr\_f))^2}{2\ln(std\_f)^2}\right) + \frac{N_{c\_c}}{\sqrt{2\pi} \ln(std\_c)} \exp\left(-\frac{(\ln(r) - \ln(nmr\_c))^2}{2\ln(std\_c)^2}\right)$$

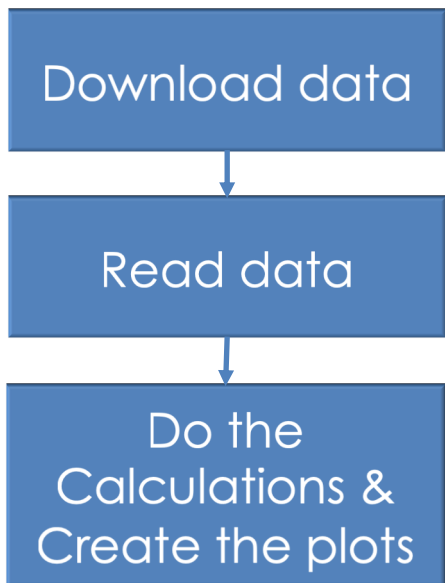
$$N_{c\_f} = \sum \left( radius\_AER\_dif \cdot \left( \frac{1}{radius\_AER\_mean} \frac{volc\_f}{\sqrt{2\pi}std\_f} \frac{\exp\left(-\frac{((\ln(radius\_AER\_mean) - \ln(vmr\_f))^2)}{2std\_f^2}\right)}{\frac{4\pi}{3}radius\_AER\_mean^3} \right) \right)$$

where

$$N_{c\_c} = \sum \left( radius\_AER\_dif \cdot \left( \frac{1}{radius\_AER\_mean} \frac{volc\_c}{\sqrt{2\pi}std\_c} \frac{\exp\left(-\frac{((\ln(radius\_AER\_mean) - \ln(vmr\_c))^2)}{2std\_c^2}\right)}{\frac{4\pi}{3}radius\_AER\_mean^3} \right) \right)$$

$N_{c\_f}$ : the total number concentration of fine mode

$N_{c\_c}$ : the total number concentration of coarse mod



## 2<sup>nd</sup> step

In order to **calculate the six parameters of the number size distribution** ( $N_{c\_f}$ ,  $N_{c\_c}$ ,  $n_{mr\_f}$ ,  $n_{mr\_c}$ ,  $std\_f$ ,  $std\_c$ ) **based on** the six parameters of the **volume size distribution** ( $volc\_f$ ,  $volc\_c$ ,  $vmr\_f$ ,  $vmr\_c$ ,  $std\_f$ ,  $std\_c$ ) we use these formulas:

$$\frac{dN}{d\ln r} = \frac{N_{c\_f}}{\sqrt{2\pi} \ln(std\_f)} \exp\left(-\frac{(\ln(r) - \ln(n_{mr\_f}))^2}{2\ln(std\_f)^2}\right) + \frac{N_{c\_c}}{\sqrt{2\pi} \ln(std\_c)} \exp\left(-\frac{(\ln(r) - \ln(n_{mr\_c}))^2}{2\ln(std\_c)^2}\right)$$

$$n_{mr\_f} = \frac{\exp(\ln(2 - vmr\_f) - 3std\_f^2)}{2}$$

where

$n_{mr\_f}$ ,  $n_{mr\_c}$ :  
the number mean radius for fine and coarse mode

$$n_{mr\_c} = \frac{\exp(\ln(2 - vmr\_c) - 3std\_c^2)}{2}$$

The **geometric standard deviation** (“ $std\_f$ ” and “ $std\_c$ ”) **does not change** for number and volume size distributions.



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Do the  
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Create the plots

## 2<sup>nd</sup> step

After calculating  $dN/d\ln r$  with the formulas presented in the previous slides, we calculate the volume size distribution using the following formula:

$$\frac{dV}{d\ln r} = \frac{dN}{d\ln r} \frac{4\pi \text{radius\_AER\_mean}^3}{3}$$

This formula uses the assumption **that the size does not change within the size bin and is equal to the mean radius of the size bin (radius\_AER\_mean)**.



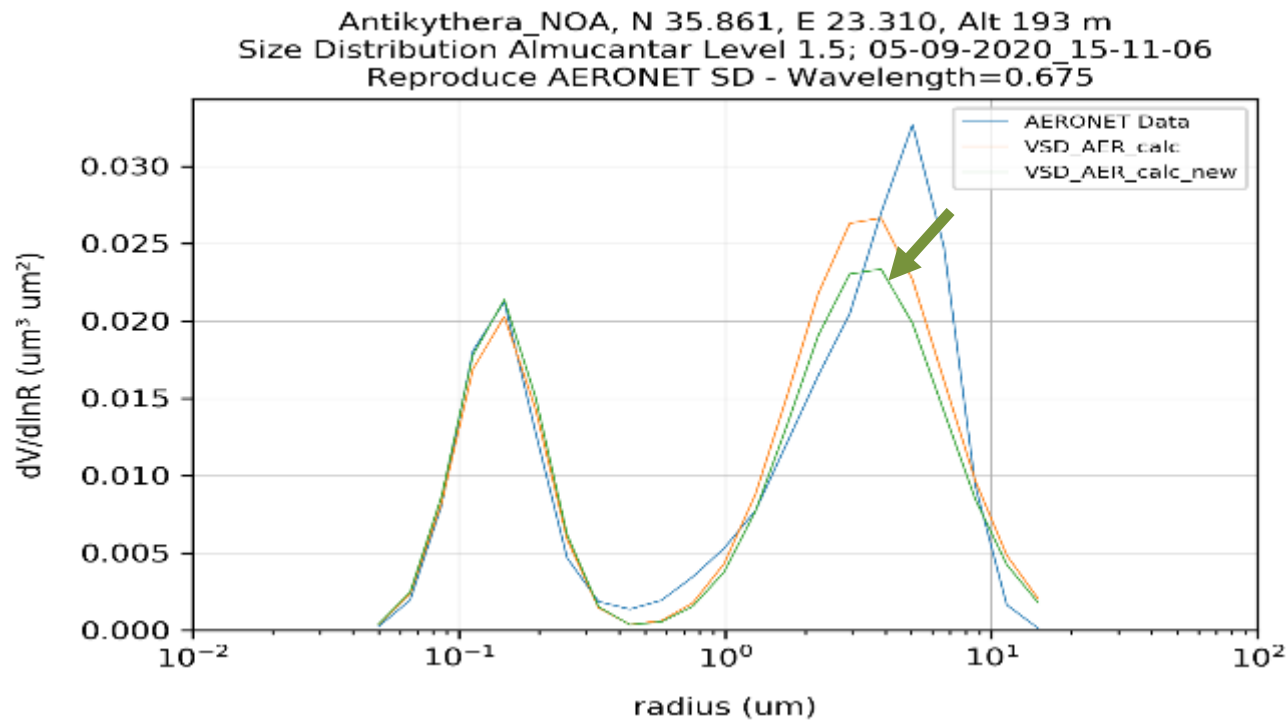
Download data

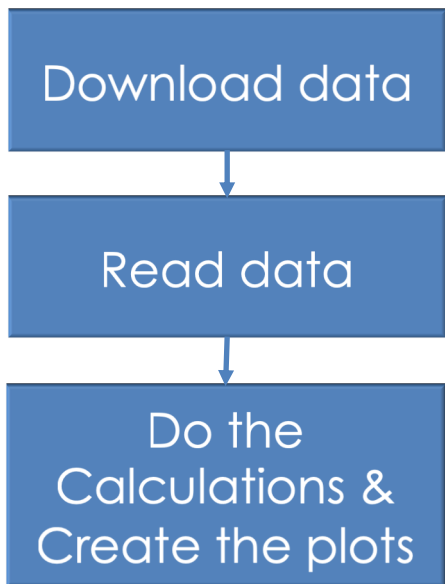
Read data

Do the  
Calculations &  
Create the plots

## 3<sup>ed</sup> step

The **green line**, is the **reproduction of the bimodal volume size distribution**, using the calculated number size distribution. This line is used to verify our calculations  $\frac{dN}{d\ln r}$ , since in the ideal case it should coincide with the orange line (i.e. the bimodal log-normal fit of the 22-size-bin volume size distribution).





## 3<sup>ed</sup> step

This does not happen all the times, mainly due to the limitation in equation to use a representative value for the radius of the size bin.

$$\frac{dV}{d\ln r} = \frac{dN}{d\ln r} \frac{4\pi \text{radius\_AER\_mean}^3}{3}$$

$$\frac{dN}{d\ln r} = \frac{Nc_f}{\sqrt{2\pi} \ln(std_f)} \exp\left(-\frac{(\ln(r) - \ln(nmr_f))^2}{2\ln(std_f)^2}\right) + \frac{Nc_c}{\sqrt{2\pi} \ln(std_c)} \exp\left(-\frac{(\ln(r) - \ln(nmr_c))^2}{2\ln(std_c)^2}\right)$$

$$Nc_f = \sum \left( \text{radius\_AER\_dif} \cdot \left( \frac{1}{\text{radius\_AER\_mean}} \frac{\text{volc}_f}{\sqrt{2\pi} \text{std}_f} \exp\left(-\frac{(\ln(\text{radius\_AER\_mean}) - \ln(\text{vnr}_f))^2}{2\text{std}_f^2}\right) \frac{4\pi}{3} \text{radius\_AER\_mean}^3 \right) \right)$$

$$Nc_c = \sum \left( \text{radius\_AER\_dif} \cdot \left( \frac{1}{\text{radius\_AER\_mean}} \frac{\text{volc}_c}{\sqrt{2\pi} \text{std}_c} \exp\left(-\frac{(\ln(\text{radius\_AER\_mean}) - \ln(\text{vnr}_c))^2}{2\text{std}_c^2}\right) \frac{4\pi}{3} \text{radius\_AER\_mean}^3 \right) \right)$$





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Read data

Do the  
Calculations &  
Create the plots

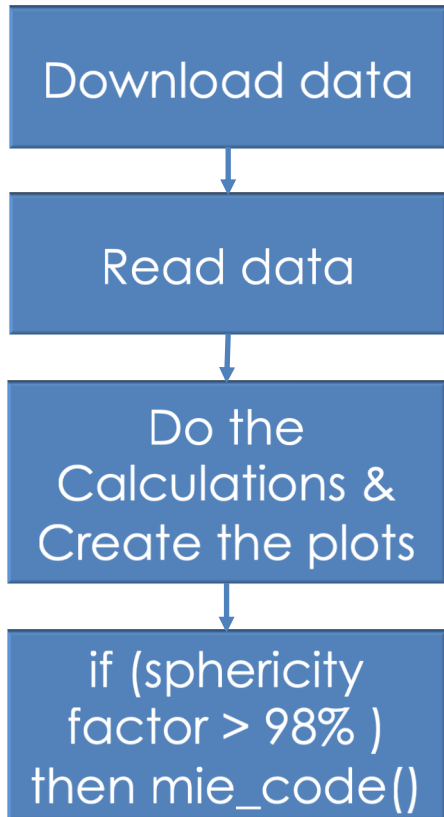
if (sphericity  
factor > 98% )  
then mie\_code()

## 4<sup>th</sup> step

The Mie Scattering code is valid only for **particles with spherical shapes**.

## More specifically

The **sphericity factor** needs to be greater than **98%**.

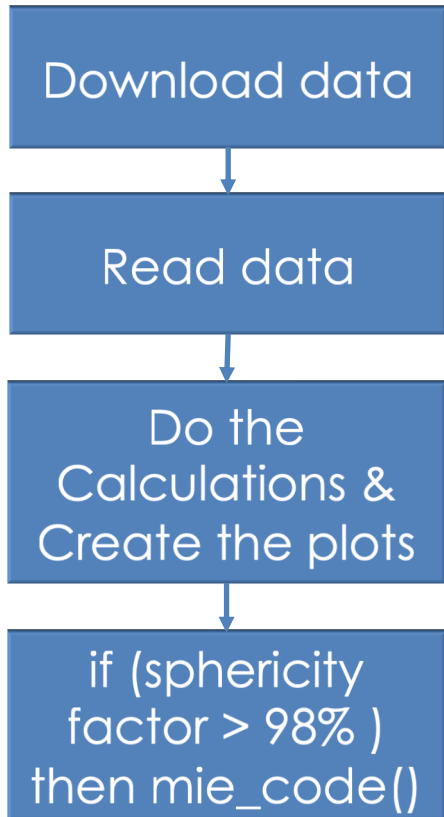


## 4<sup>th</sup> step

### Mie scattering code

#### Inputs

- The user-defined wavelength
- The parameters of the number size distribution ( $Nc_f$ ,  $Nc_c$ ,  $nmr_f$ ,  $nmr_c$ ,  $std_f$  and  $std_c$ )
- The real and imaginary part of the refractive index (considered to be the same for fine and coarse particles)



## 4<sup>th</sup> step

The `mie_code()` function **calculates**:

- The **number concentration in each sub-bin** for fine and coarse mode respectively based on the following formula:

$$dcon\_c = Nc\_c \exp\left(-1 \left(\ln \frac{r}{nmr\_c}\right)^2 / (2 std\_c^2)\right) \frac{1}{r std\_c \sqrt{2\pi}}$$

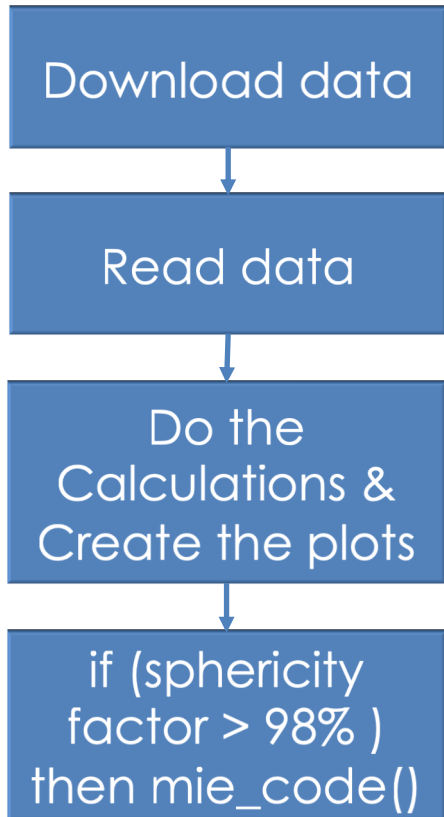
where dcon\_c: the number concentration in each bin for coarse mode

$Nc\_c$ : the total number concentration of coarse mode

$nmr\_c$ : number geometric mean radius of coarse mode

$r$ : bin center radii

$std\_c$ : geometric standard deviation of coarse mode



## 4<sup>th</sup> step

$$dcon_c = Nc_c \exp\left(-1 \left(\ln \frac{r}{nmr_c}\right)^2 / (2 std_c^2)\right) \frac{1}{r std_c \sqrt{2\pi}}$$

The mie\_code() function **calculates**:

- The **volume size distribution** ( $SD_{dv}dlnr$ ) for fine and coarse mode based on the following formula:

$$SD_{dv}dlnr = r \left(\frac{4\pi}{3}\right) r^3 dcon$$

and a **fourth line** is created in the plot.



Download data

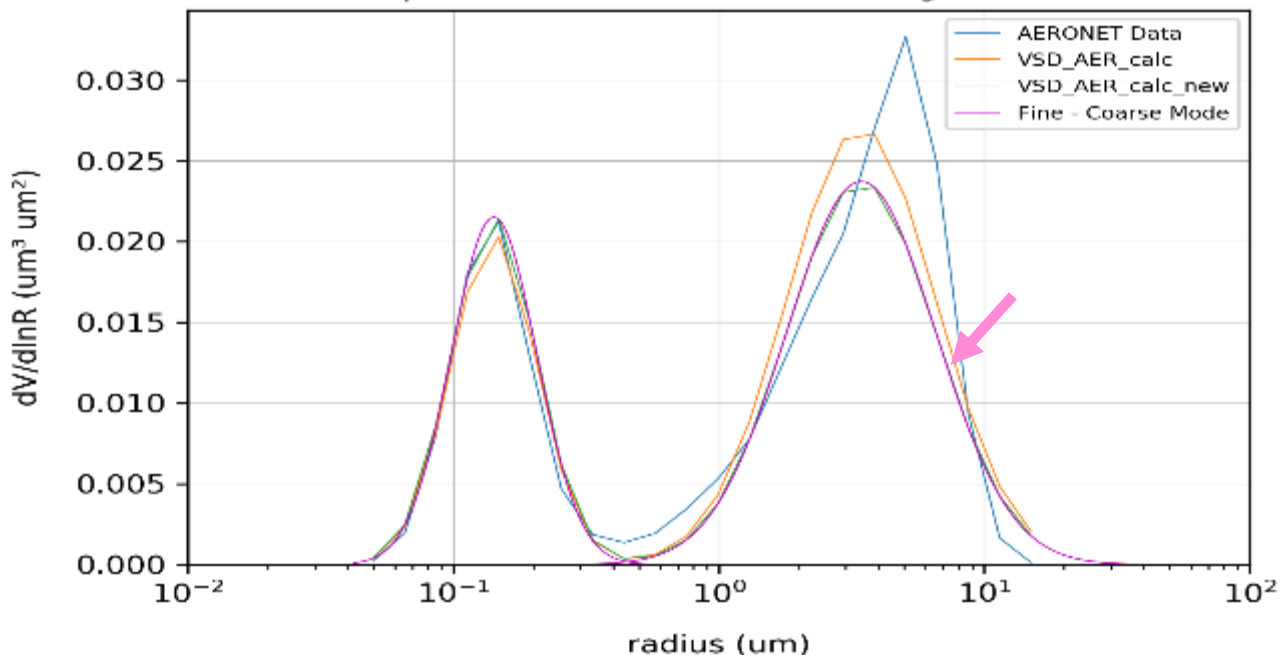
Read data

Do the  
Calculations &  
Create the plots

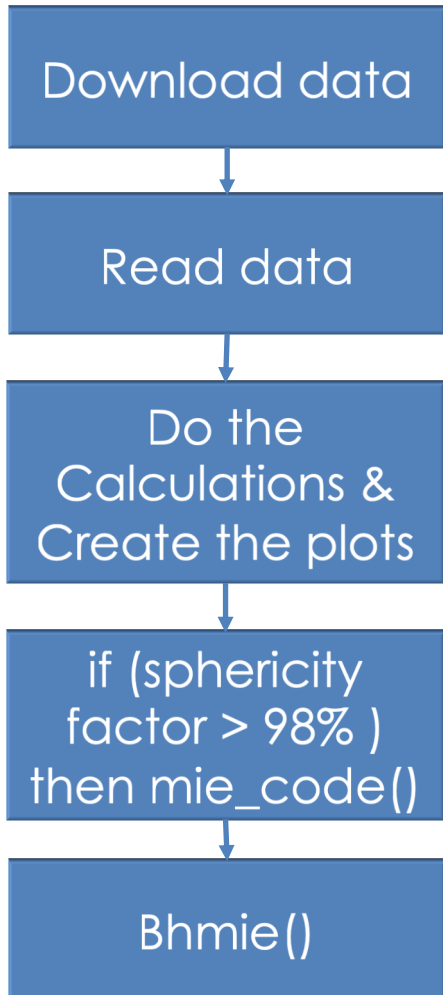
if (sphericity  
factor > 98% )  
then mie\_code()

## 4<sup>th</sup> step

Antikythera\_NOA, N 35.861, E 23.310, Alt 193 m  
Size Distribution Almucantar Level 1.5; 05-09-2020\_15-11-06  
Reproduce AERONET SD - Wavelength=0.675



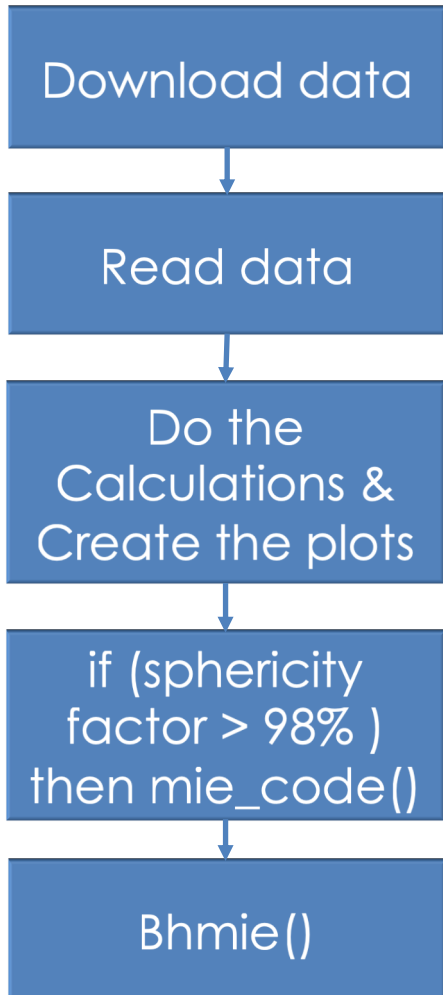
The **pink line** is a **verification** point in the algorithm (it should coincide with the green line), showing that the size distribution we provide in the input is indeed the one that is used in the Mie calculations.



## 4<sup>th</sup> step

The `mie_code()` function uses **the `bhmie()` function** that calculates the **optical properties** of the **individual particle**, with specific size and refractive index.

The `bhmie()` function that was used in this thesis, was written in 1983 from **Craig F. Bohren and Donald R. Huffman**.

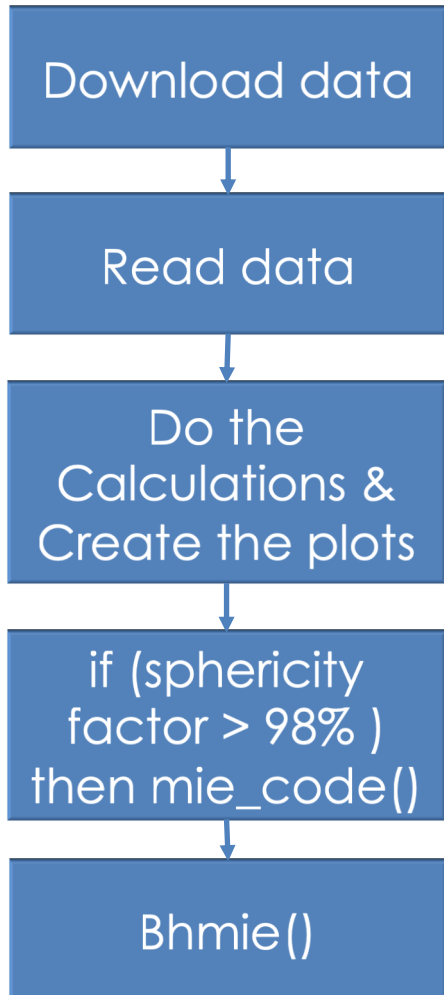


## 4<sup>th</sup> step

The `mie_code()` function uses **the `bhmie()` function** that calculates the **optical properties** of the **individual particle**, with specific size and refractive index.

## Inputs

- The size parameter
- The refractive index
- The number of scattering angles



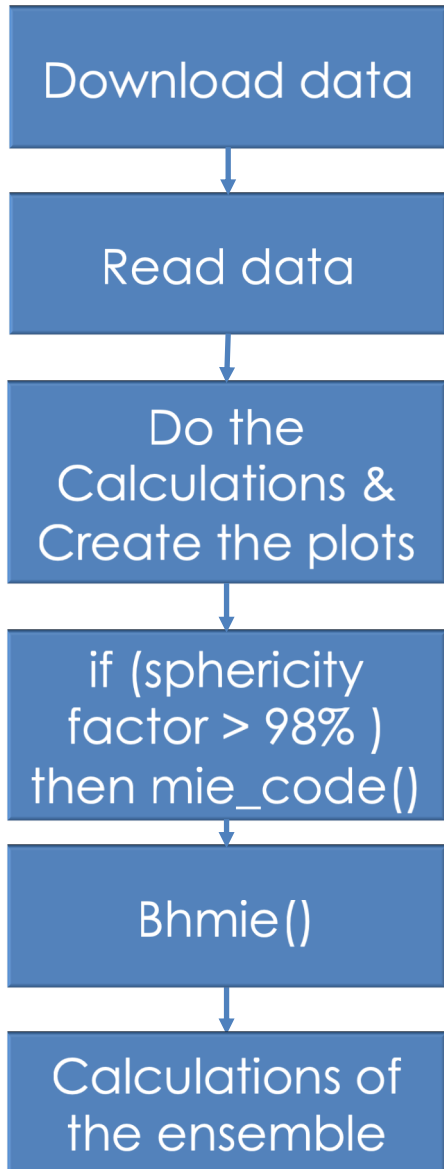
## 4<sup>th</sup> step

The `mie_code()` function uses the `bhmie()` function that calculates the **optical properties** of the **individual particle**, with specific size and refractive index.

## Outputs

- The **extinction efficiency**
- The **scattering efficiency**
- The **backscatter efficiency**
- The **absorption efficiency**





## 5<sup>th</sup> step

In this step we calculate the **optical properties** of the **particle ensemble**, taking into account the size distribution of the particles in the ensemble:

### ❑ The extinction coefficient

$$\varepsilon_{ext} = \int N(r)\sigma_{ext} dr = \int N(r)Q_{ext} \pi r^2 dr$$

### ❑ The scattering coefficient

$$\varepsilon_{sca} = \int N(r)\sigma_{sca} dr = \int N(r)Q_{sca} \pi r^2 dr$$

### ❑ The backscatter coefficient

$$\varepsilon_{bck} = \int N(r)Q_{bck} \left(\frac{\lambda}{\pi r}\right)^2 dr$$

### ❑ The absorption coefficient

$$\varepsilon_{abs} = \int N(r)\sigma_{abs} dr = \int N(r)Q_{abs} \pi r^2 dr$$

# Results of the study



On the date of 27 of April on 2020, there are **no** available **data** from the Antikythera AERONET station, and the user is informed with the following message:

```
Start date is: year=2019&month=4&day=27
End date is: year2=2019&month2=4&day2=27&hour2=23
Wavelength is: 0.675
https://aeronet.gsfc.nasa.gov/cgi-bin/print_web_data_inv_v3?
site=Antikythera_NOA&AVG=10&ALM15=1&if_no_html=1&product=SIZ&year=201
9&month=4&day=27&year2=2019&month2=4&day2=27&hour2=23
https://aeronet.gsfc.nasa.gov/cgi-bin/print_web_data_inv_v3?
site=Antikythera_NOA&AVG=10&ALM15=1&if_no_html=1&product=VOL&year=201
9&month=4&day=27&year2=2019&month2=4&day2=27&hour2=23
https://aeronet.gsfc.nasa.gov/cgi-bin/print_web_data_inv_v3?
site=Antikythera_NOA&AVG=10&ALM15=1&if_no_html=1&product=ASY&year=201
9&month=4&day=27&year2=2019&month2=4&day2=27&hour2=23
https://aeronet.gsfc.nasa.gov/cgi-bin/print_web_data_inv_v3?
site=Antikythera_NOA&AVG=10&ALM15=1&if_no_html=1&product=RIN&year=201
9&month=4&day=27&year2=2019&month2=4&day2=27&hour2=23
Site=Antikythera_NO no data is available in this period: starting
date=27/4/2019 and ending date=27/4/2019
```



In this case, the **start and end date** are the **same**. The day is the 18th of July on 2019 and the **wavelength of refractive index is equal to 0.675**  $\mu\text{m}$ . The data for this date is available as shown in the message below:

Start date is: year=2019&month=7&day=18

End date is: year2=2019&month2=7&day2=18

Wavelength is: 0.675

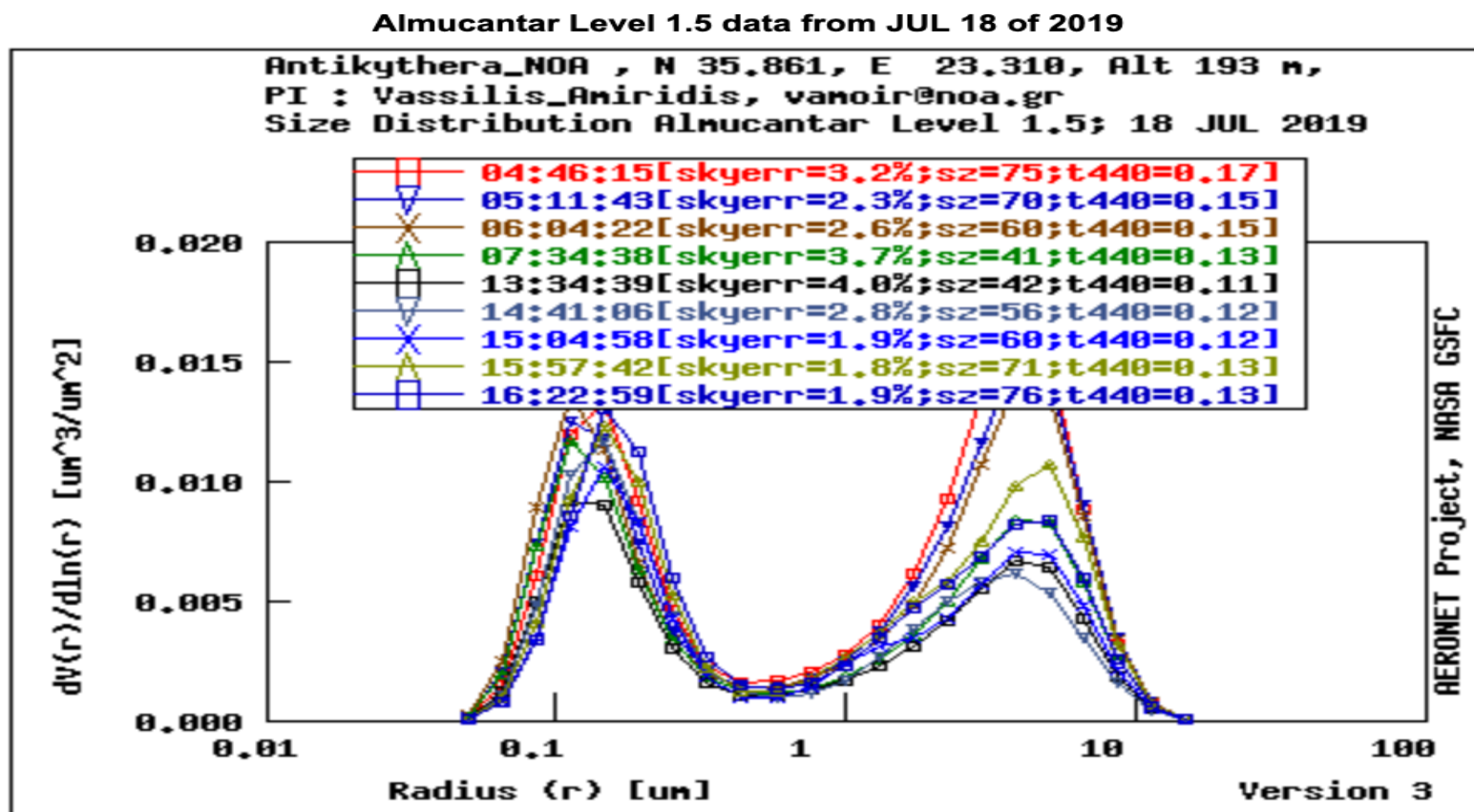
```
https://aeronet.gsfc.nasa.gov/cgi-bin/print_web_data_inv_v3?
site=Antikythera_NOA&AVG=10&ALM15=1&if_no_html=1&product=SIZ&year=2019&month=7&da
y=18&year2=2019&month2=7&day2=18
https://aeronet.gsfc.nasa.gov/cgi-bin/print_web_data_inv_v3?
site=Antikythera_NOA&AVG=10&ALM15=1&if_no_html=1&product=VOL&year=2019&month=7&da
y=18&year2=2019&month2=7&day2=18
https://aeronet.gsfc.nasa.gov/cgi-bin/print_web_data_inv_v3?
site=Antikythera_NOA&AVG=10&ALM15=1&if_no_html=1&product=ASY&year=2019&month=7&da
y=18&year2=2019&month2=7&day2=18
https://aeronet.gsfc.nasa.gov/cgi-bin/print_web_data_inv_v3?
site=Antikythera_NOA&AVG=10&ALM15=1&if_no_html=1&product=RIN&year=2019&month=7&da
y=18&year2=2019&month2=7&day2=18
```



# Results of 18/07/2019



This certain day, based on the AERONET site, there are **nine different retrievals** of the volume size distribution for different times during the day, as shown in the figure below: at 04:46:15, 05:11:43, 06:04:22, 07:34:38, 13:34:39, 14:41:06, 15:04:58, 15:57:42, 16:22:59 UTC





**Time: 04:46:15**

```
Creating plot for SIZ_Antikythera_NO_18-07-2019_04-46-15
SIZ_Antikythera_NO_18-07-2019_04-46-15 Sphericity_Factor 99.00 %
volc_f=0.014000 vmr_f=0.148000 std_f=0.413000 volc_c=0.023000 vmr_c=3.132000
std_c=0.656000 rirp440=1.600000 rirp675=1.541400 rirp870=1.521200
rirp1020=1.524100 riip440=0.007420 riip675=0.006930 riip870=0.006592
rirp1020=0.006531
RI 0.675
```

The **AERONET product for this case** provides the following aerosol properties:

- ❑ **Total volume concentration** of fine mode (volc\_f) : 0.014000 cm<sup>3</sup> cm<sup>-3</sup>
- ❑ **Geometric mean volume radius** of fine mode (vmr\_f): 0.148000 μm
- ❑ **Geometric standard deviation** of fine mode (std\_f): 0.413000
- ❑ **Total volume concentration** of coarse mode (volc\_c): 0.023000 cm<sup>3</sup> cm<sup>-3</sup>
- ❑ **Geometric mean volume radius** of coarse mode (vmr\_c): 3.132000 μm
- ❑ **Geometric standard deviation** of coarse mode (std\_c): 0.656000



**Time: 04:46:15**

```
Creating plot for SIZ_Antikythera_NO_18-07-2019_04-46-15
SIZ_Antikythera_NO_18-07-2019_04-46-15 Sphericity_Factor 99.00 %
volc_f=0.014000 vmr_f=0.148000 std_f=0.413000 volc_c=0.023000 vmr_c=3.132000
std_c=0.656000 rirp440=1.600000 rirp675=1.541400 rirp870=1.521200
rirp1020=1.524100 riip440=0.007420 riip675=0.006930 riip870=0.006592
riip1020=0.006531
RI 0.675
```

The **AERONET product for this case** provides the following aerosol properties:

- ❑ **Real part of the refractive index** at 440 nm (rirp440): 1.600000
- ❑ **Real part of the refractive index** at 675 nm (rirp675): 1.541400
- ❑ **Real part of the refractive index** at 870 nm (rirp870): 1.521200
- ❑ **Real part of the refractive index** at 1020 nm (rirp1020): 1.524100
  
- ❑ **Imaginary part of the refractive index** at 440 nm (riip440): 0.007420
- ❑ **Imaginary part of the refractive index** at 675 nm (riip675): 0.006930
- ❑ **Imaginary part of the refractive index** at 870 nm (riip870): 0.006592
- ❑ **Imaginary part of the refractive index** at 1020 nm (riip1020): 0.006531



**Time: 04:46:15**

ext\_all=7.909338890331058e-05 sca\_all=7.346999408914502e-05  
abs\_all=5.623394814165551e-06 bac\_all=2.032112840574509e-06  
ext\_f=6.52076043790471e-05 sca\_f=6.241395193588643e-05  
abs\_f=2.7936524431606673e-06 bac\_f=1.3391959966208895e-06  
ext\_c=1.388578452426348e-05 sca\_c=1.1056042153258596e-05  
abs\_c=2.829742371004884e-06 bac\_c=6.929168439536193e-07

The calculated **outputs** for this case are the following:

- Extinction coefficient of the ensemble** (ext\_all): 7.909338 Km<sup>-1</sup>
- Scattering coefficient of the ensemble** (sca\_all): 7.346999 Km<sup>-1</sup>
- Absorption coefficient of the ensemble** (abs\_all): 5.623394 Km<sup>-1</sup>
- Backscatter coefficient of the ensemble** (bac\_all): 2.032112 Km<sup>-1</sup> sr<sup>-1</sup>
  
- Extinction coefficient of the fine mode** (ext\_f): 6.520760 Km<sup>-1</sup>
- Scattering coefficient of the fine mode** (sca\_f): 6.241395 Km<sup>-1</sup>
- Absorption coefficient of the fine mode** (abs\_f): 2.793652 Km<sup>-1</sup>
- Backscatter coefficient of the fine mode** (bac\_f): 1.339195 Km<sup>-1</sup> sr<sup>-1</sup>





**Time: 04:46:15**

ext\_all=7.909338890331058e-05 sca\_all=7.346999408914502e-05  
abs\_all=5.623394814165551e-06 bac\_all=2.032112840574509e-06  
ext\_f=6.52076043790471e-05 sca\_f=6.241395193588643e-05  
abs\_f=2.7936524431606673e-06 bac\_f=1.3391959966208895e-06  
ext\_c=1.388578452426348e-05 sca\_c=1.1056042153258596e-05  
abs\_c=2.829742371004884e-06 bac\_c=6.929168439536193e-07

The calculated **outputs** for this case are the following:

- ❑ **Extinction coefficient of the coarse mode** (ext\_c): 1.388578 Km<sup>-1</sup>
- ❑ **Scattering coefficient of the coarse mode** (sca\_c): 1.105604 Km<sup>-1</sup>
- ❑ **Absorption coefficient of the coarse mode** (abs\_c): 2.829742 Km<sup>-1</sup>
- ❑ **Backscatter coefficient of the coarse mode** (bac\_c): 6.929168 Km<sup>-1</sup> sr<sup>-1</sup>

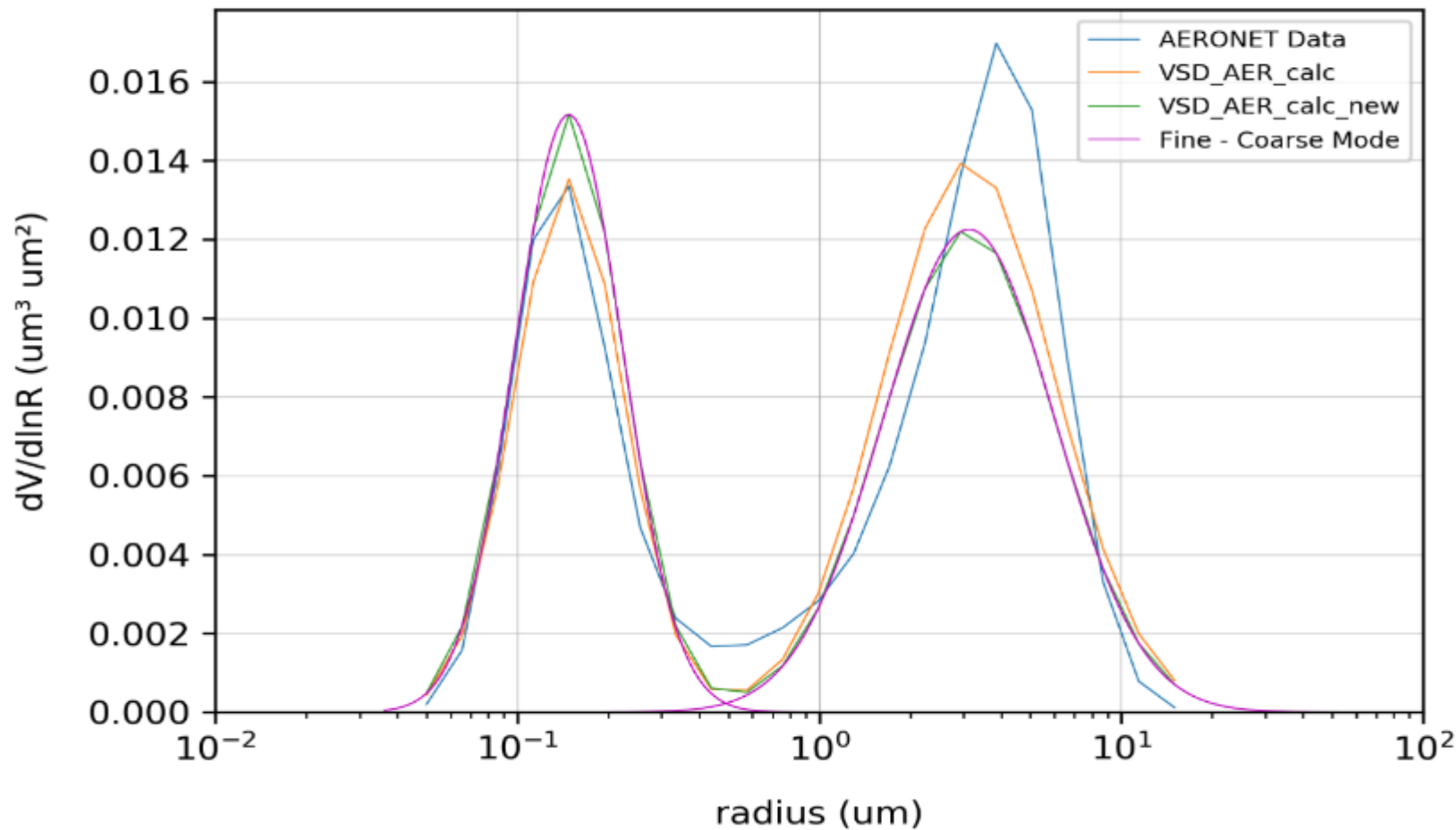


# Results of 18/07/2019



Time: 04:46:15

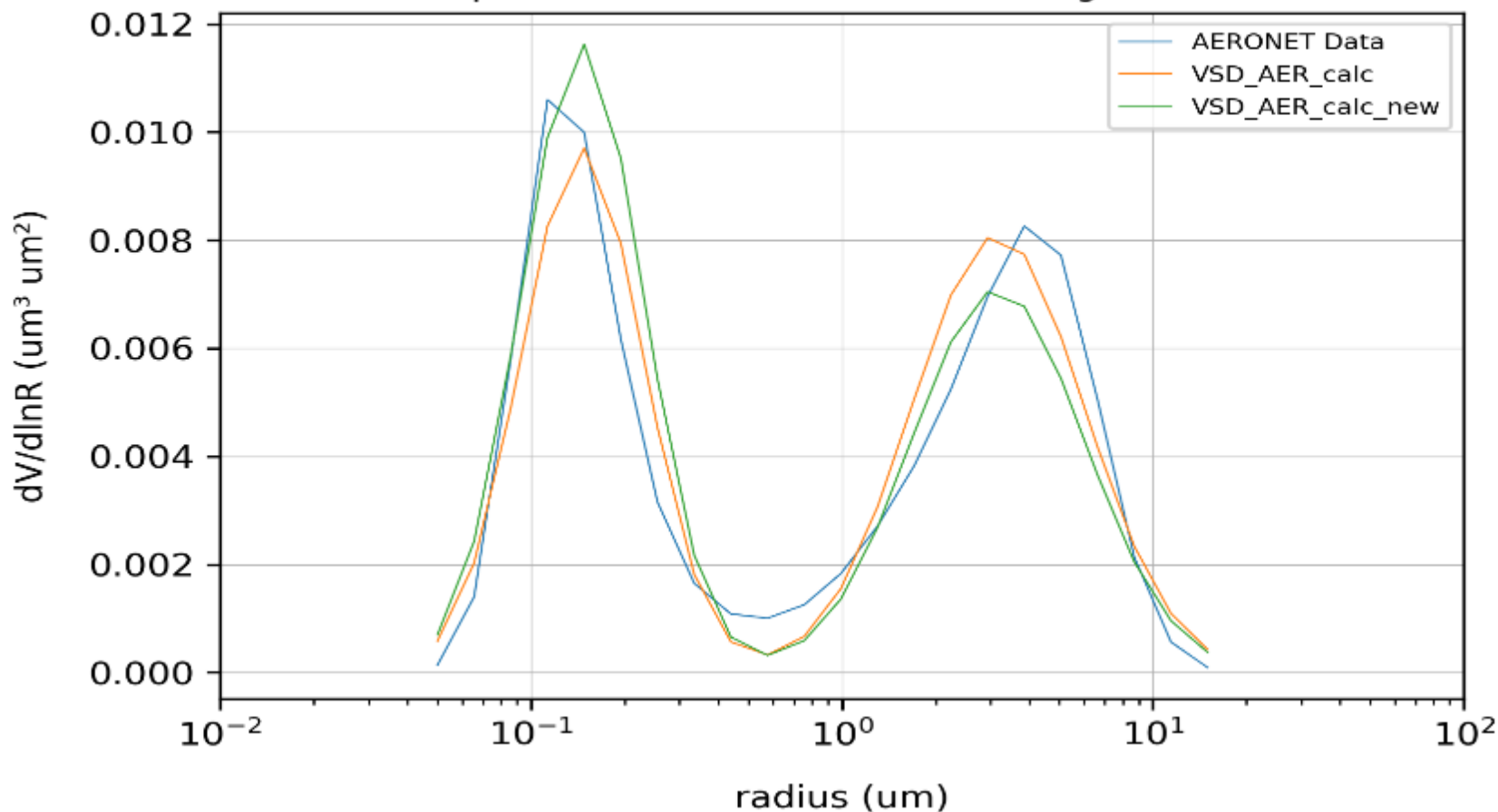
Antikythera\_NOA, N 35.861, E 23.310, Alt 193 m  
Size Distribution Ålmuçantar Level 1.5; 18-07-2019\_04-46-15  
Reproduce AERONET SD - Wavelength=0.675





Time: 07:34:38

Antikythera\_NOA, N 35.861, E 23.310, Alt 193 m  
Size Distribution Almucantar Level 1.5; 18-07-2019\_07-34-38  
Reproduce AERONET SD - Wavelength=0.675



Future work



The developed tool focuses mainly on the lidar-related optical properties.



**These properties can be used in the future step of this work, for combining the sun-photometer with lidar measurements for better characterization of aerosol vertical properties.**

# Summary



**SUMMARY**

The goal of this work was to create an **automated tool** that calculates in near **real-time the optical properties** of aerosol particles observed in the **PANGEA observatory**, when the particles have **spherical shapes**. For this scope we use the **sun-photometer measurements** in the station, the **AERONET inversion products** and the **Mie scattering code**.



# Thanks for your attention!



- ▣ My supervisor **Dr. Vassilis Amiridis**
- ▣ The postdoctoral researcher **Alexandra Tsekeri**





# Questions ?



**Let's try it together...**

AERONET

Aerosols

Mie code

Remote Sensing



PANGEA

optical properties

microphysical properties

Backup Slides



AERONET data display uses a **user-friendly interface** that includes **graphical representations** to display the optical and microphysical properties of the aerosol particles.

There are **three versions** available for the **AERONET aerosol product**:

- ▣ Version 1.0
- ▣ Version 2.0
- ▣ Version 3.0

## Data quality levels for Direct Sun algorithm:

- ▣ Level 1 (unscreened)
- ▣ Level 1.5 (cloud-screened and quality controlled)
- ▣ Level 2.0 (quality-assured)

## Data quality levels for Direct Sun and Inversion algorithm:

- ▣ Inversion Level 1.5 (cloud-screened and quality controlled)
- ▣ Inversion Level 2.0 (quality-assured)



Time: 05:58:11

Antikythera\_NOA, N 35.861, E 23.310, Alt 193 m  
Size Distribution  $\bar{A}$ Imucantar Level 1.5; 30-06-2020\_05-58-11  
Reproduce AERONET SD - Wavelength=0.44

